

Highly-Excited State Properties of Cumulenone Chlorides in the Vacuum-Ultraviolet[†]

Quynh L. D. Nguyen,^{*a} William K. Peters,^{a‡} and Ryan C. Fortenberry^b

Recent observations of chloromethane in interstellar environments suggest that other organohalogens, which are known to be critically important in Earth's atmosphere, may also be of significance beyond our own terrestrial veil. This raises the question of how such molecules behave under extreme conditions such as when exposed to vacuum ultraviolet (VUV) radiation. VUV photons promote molecules to highly excited states that fragment in non-statistical patterns controlled by the initial femtosecond dynamics. A detailed understanding of VUV-driven photochemistry in complex organic molecules that consist of more than one functional group is a particularly challenging task. This quantum chemical analysis reports the electronic states and ionization potentials up to the VUV range (6 - 11 eV) of the chlorine-substituted cumulenone series molecules. The valence and Rydberg properties of lone-pair terminated, π -conjugated systems are explored for their potential resonance with lone pairs from elsewhere in the system. The carbon chain elongation within the family ClHC_nO, where $n=1-4$, influences the electronic excitations, associated wavefunctions, and ionization potentials of the molecules. The predicted geometries and ionization potentials are in good agreement with the available experimental photoelectron spectra for formyl chloride and chloroketene, $n=1-2$. Furthermore, comparison between the regular cumulenone species and the corresponding chlorinated derivatives exhibit similar behaviors especially for $n=3$, where the allene backbone in propadienone chloride is severely bent. Most notably for the excited states is that the Rydberg character becomes more dominant as the energy increases, with some retaining valence characters.

1 Introduction

Most small molecules concentrate their oscillator strength in the vacuum ultraviolet (VUV)¹, and a detailed knowledge of the corresponding excited states is critical to understanding photophysics and photochemistry in environments containing both appreciable VUV photon flux and significant molecular density. Such environments include circumstellar dust clouds, the upper portions of planetary atmospheres, and the immediate vicinity of high-energy transient phenomena such as lightning strikes or explosions²⁻⁴. For example, on our own planet, incident VUV flux interacting with molecular flux from below influences the ionosphere-thermosphere coupling^{5,6}. Additionally, in exoplanetary atmospheres, VUV-driven photochemistry can produce nonequilibrium chemical compositions that may otherwise be interpreted as biosignatures⁷. In protoplanetary disks, a deeper understanding of how complex molecules form in the presence of an overwhelming VUV photon flux is of critical importance to modeling the chemical composition of nascent planetary atmospheres^{8,9}.

In small polyatomic molecules, the energy range from about 5 or 6 eV up to the ionization threshold, roughly 10 eV, presents particular challenges both experimentally and theoretically^{10,11}.

In this energy range, molecules display both highly-excited valence states and low-lying Rydberg states, often strongly mixed with each other. These states generally lead to a cascade of non-statistical fragmentation with branching ratios dominated by early-time (femtosecond) dynamics. Detailed knowledge is hard to come by, however, as large densities of states and highly diffuse wavefunctions make *ab initio* calculations prohibitively expensive. Experimentally, the same large densities of states combine with short lifetimes to cause congested spectra that require nonlinear spectroscopic techniques to interrogate, but convenient and flexible laser sources have only recently become available.

Previous results, both experimental and theoretical, on initial femtosecond dynamics and their influence on the relaxation of VUV excited states have pointed toward a critical role played by Rydberg–valence mixing. Experiment and high level calculations on ethylene has indicated that the $\pi^* \leftarrow \pi$ quickly transfers to the $3s$ –Rydberg state within less than 10 fs through ultrafast internal conversion and then transfers back to the valence state within 60 fs^{12,13}. Ultrafast photoelectron spectroscopy on the nitrogen-containing heterocycles pyrrole, N-methyl pyrrole, and 2,5-dimethyl pyrrole have indicated the $3p$ –Rydberg state is populated for tens of femtoseconds before Rydberg–valence mixing allows access to conical intersections that drive electronic relaxation^{14,15}. Ultrafast photoelectron-photoion coincidence measurements combined with high-level calculations have revealed strong nonadiabatic behavior as well as Rydberg–valence mixing both play key roles in highly excited states of methyl azide, further emphasizing the difficulty of such studies¹⁶.

As more complex molecules are considered, there is an open question of what may happen when multiple Rydberg series built

^a Department of Chemistry, University of Colorado Boulder. E-mail: Quynh.L.Nguyen@colorado.edu

^b Department of Chemistry & Biochemistry, University of Mississippi, University, MS 38677-1848, U.S.A.

[†] Electronic Supplementary Information (ESI) available: Additional details of calculations and cartesian coordinates. . See DOI: 10.1039/cXCP00000x/

[‡] Present address: Los Alamos National Laboratory

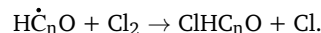
on different ionization centers are present. For example, in an organic molecule with two functional groups present, Rydberg series converging to ionization from the two different groups may both exist. Interestingly, a near-degeneracy between low- n Rydberg states of different series would be quite common. Using typical quantum defects [δ is about 1 for an s -wave and δ is about 0.5 for a p -wave^{10,11}], we would expect HOMO $\rightarrow 3p$ and (HOMO-1) $\rightarrow 3s$ to be within a few tenths of an electronvolt anytime the first two ionization potentials are separated by 1-1.5 eV, which is a reasonably common situation for organic molecules¹⁷. In the presence of these near-degeneracies, we may expect significant mixing between different Rydberg series, either by directly coupling to each other (similar to hole-mixing or non-Koopman's effects in photoelectron spectra) or by both coupling to the same valence state. Indeed, exactly this sort of situation was observed in methyl azide¹⁶ where the excited state consisted of two diabatic Rydberg states and a diabatic valence state ($\pi^{nb} \rightarrow 3p$, $\sigma_{CN} \rightarrow 3s$ and $\sigma_{CN} \rightarrow \sigma^*$). Such mixing has been noted in a few molecules, including CO, O₂, N₂, and HN₃, but the phenomenon has not been extensively explored¹⁸⁻²¹.

Formyl chloride (ClCHO)—a derivative of formaldehyde with a hydrogen replaced by chlorine—presents a potential prototype for VUV photophysics in the presence of two Rydberg series. It contains two distinguishable ionization centers—lone pairs on the oxygen atom and those on the chlorine atom—that appear relatively localized in the photoelectron spectra²² but whose Rydberg states are likely to appear in near-degeneracies. Furthermore, the species is known to play an important role in terrestrial photochemistry²³⁻²⁷ and due to recent observations of chloroalkanes in the interstellar medium²⁸⁻³² may be expected to exist in a variety of astrochemical settings, indicating that its VUV excited states are not merely of theoretical interest.

In Earth's atmosphere, ClHCO is a relatively transient molecule that has been the subject of many high resolution infrared and microwave spectroscopic studies due to its significant contributions to tropospheric and stratospheric pollution²³. It is typically observed under tropospheric conditions as a major product of reactions between chlorinated species with volatile organic compounds—*isoprene* with Cl atoms²⁴, *H₂O* radical with CHCl₂O^{25,26}, and *OH* radical with vinyl chloride or trichloroethene²⁷. Additionally, formyl chloride is either a major product or an intermediate in atmospheric breakdowns of Cl-bearing hydrocarbons that are, in turn, major players in the degradation of the ozone layer. From here, theory predicts that ClCHO can dissociate via Cl-C bond cleavage via the internal conversion or intersystem crossing from the lowest singlet and triplet states³³. It may also react with H radicals to either produce Cl $\dot{C}O$ + H₂ or Cl $\dot{C}O$ + HCl, where the hydrogen abstraction channel dominates²³. Some of these are among the most dominant and reactive species in the atmosphere, and thus, understanding the photophysical properties of formal chloride could potentially yield insights in pollutant mitigation within the ozone layer.

Considering the abundance of extended linear chains that have been detected in interstellar and circumstellar environments³⁴⁻³⁷, we extend our analysis from formyl chloride to longer cumulenone chlorides. Although these species have

not yet been detected, the non-chlorinated cumulenone ketone species—formaldehyde (H₂CO), ketene (H₂C=C=O), propadienone (H₂C=C=C=O), and their corresponding radicals post hydrogen cleavage (H \dot{C}_nO)—have been detected in the atmosphere^{38,39} and interstellar clouds⁴⁰⁻⁴². The dehalogenated radicals of these species have also been discovered in laboratory settings with some also observed in the ISM, suggesting that reactions similar to those of formyl chloride are certainly possible in similar conditions^{43,44}. These cumulene neutrals and radicals are bathed in a chlorine environment, which may eventually lead to the formation of chlorinated cumulenones via



Following the same scheme, ClHCO and Cl have been produced exclusively with no byproducts, where formaldehyde radicals, H $\dot{C}O$, were generated via photolysis^{43,45,46}.

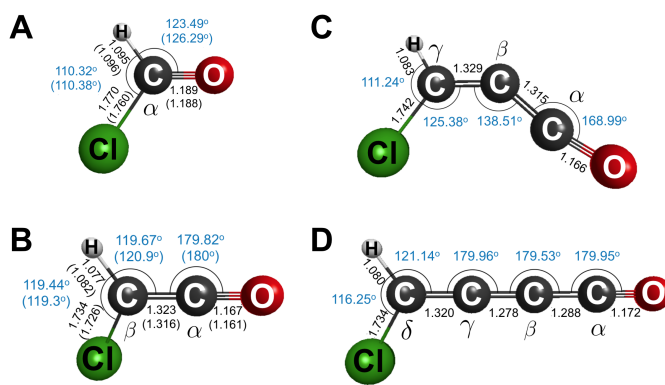


Fig. 1 Calculated neutral ground state geometries of cumulenone chlorides: Geometry optimizations are done at CCSD(T)/d-aug-cc-pV(T+d)Z for (A) formyl chloride, (B) chloroketene, (C) propadienone chloride, and (D) butatrienone chloride. Corresponding microwave spectroscopic measurements are shown in parentheses for (A) and (B).^{47,48}

Beyond our own atmosphere on Earth, chlorinated organics have been observed on Mars by the *Viking* and *Curiosity* landers^{28,29}, and also the interstellar medium (ISM)³⁰⁻³². Particularly, haloalkanes have recently been detected in star-forming regions where the UV photon flux is relatively high. Most notably, chloromethane recently became the first Cl-bearing organohalogen reported in the triple protostar system IRAS 16293-2422 where photochemistry is highly likely to influence the surrounding chemistry⁴⁹. In environments with both chloroalkanes and oxidizing species, such as Martian soil, we may suspect similar chemistry as that observed in Earth's atmosphere, including the possible presence of formyl chloride^{28,29,50}.

In this paper, we calculate ground state geometry, vertical and adiabatic ionization energies, and vertical excitation energies including in the VUV region for formyl chloride. We extend the analysis on ClCHO to its related species with elongated allene chains: chloroketene (ClHC=C=O, ClHC₂O), propadienone chloride (ClHC=C=C=O, ClHC₃O), and butatrienone chloride (ClHC=C=C=C=O, ClHC₄O). We compare our predicted geometries and excited state properties to experimental data in cases where measurements are available. Our findings represent an

Table 1 Theoretical bond lengths (Å) and angles (°) of the cumulenone chlorides optimized at CCSD(T)/d-aug-cc-pV(T+d)Z. The corresponding microwave experimental measurements^{47,48} are reported in parentheses underneath.

	$r_{C\text{Cl}}$	r_{CH}	r_{CO}	$r_{C_{\alpha\beta}}$	$r_{C_{\beta\gamma}}$	$r_{C_{\gamma\delta}}$	$\theta_{C\text{ClC}/C\text{ICO}}$	$\theta_{H\text{CCL}}$	θ_{OCC}	$\theta_{C_{\alpha\beta\gamma}}$
ClHCO	1.770 (1.760)	1.095 (1.096)	1.189 (1.188)				123.490 (126.29)	110.318 (110.38)		
ClHC ₂ O	1.734 (1.726)	1.077 (1.082)	1.167 (1.161)	1.323 (1.316)			119.674 (120.9)	119.436 (119.3)	179.819 (180)	
ClHC ₃ O	1.742	1.083	1.166	1.315	1.329		125.382	111.244	168.992	138.508
ClHC ₄ O	1.734	1.080	1.172	1.288	1.278	1.320	121.142	116.253	179.948	179.530

important step toward understanding the interaction between different functional groups in VUV-induced photophysics and photochemistry of complex molecules, which is already strikingly non-statistical even in simpler molecules.

2 Computational Methods

The CFOUR⁵¹ and PSI4⁵² computational packages are employed for all calculations with several correlation consistent basis sets⁵³. The ground state neutral geometries are optimized through density functional theory (DFT) with the B3LYP functionals^{54,55}, second-order Moller-Plesset perturbation theory (MP2)⁵⁶, and coupled cluster theory at the singles, doubles, and perturbative triples (CCSD(T))⁵⁷ levels of theory with the d-aug-cc-pV(T+d)Z⁵⁸ basis set. All minima are confirmed with harmonic vibrational analysis at the corresponding levels of theory, which yield all real frequencies. ClHC₄O geometry optimized at the CCSD(T) level, however, yields one imaginary frequency that will be discussed later. Vertical (VIE) and adiabatic (AIE) ionization energy calculations of all species are computed via equation of motion - ionization potential - coupled cluster singles and doubles (EOMIP-CCSD)^{59–61} with the aug-cc-pVTZ and d-aug-cc-pVTZ basis functions. The AIEs for ClHCO, ClHC₂O, and ClHC₃O are also computed with the d-aug-cc-pV(T+d)Z basis.

Vertical excitation energies are calculated at the EOM-CCSD^{62,63} level with multiple basis sets, including correlation consistent and Ahlrich's basis sets^{64,65}. Extra diffuse functions are added to the d-aug-cc-pV(T+d)Z and t-aug-cc-pV(T+d)Z implemented in PSI4 in an even-tempered manner used previously for excessively diffuse basis sets of anions^{66–69}. These bases are used for all single point calculations to fully account for the diffuseness of the electron density in potentially highly-diffuse states around the molecule. Additionally, since the computational cost scales up with the size of the molecule, a set of 6s, 6p, 2d diffuse functions located about the center-of-mass are utilized in conjunction with the aug-cc-pVDZ and aug-cc-pVTZ basis sets to compute the excited state properties of all four systems^{68,70,71} and are written in shorthand as DZ+sp_d and TZ+sp_d.

In order to characterize the excited states, a MATLAB script is written to take linear combinations of virtual t-aug-cc-pV(T+d)Z molecular orbitals that participate in the transition with the coefficients given by EOM-CCSD calculations. A python script is used in complement with VMD1.9.3 to visualize the transition orbitals for each excited state^{52,72}. A complete set of excited state orbitals for each molecule can be found in the supplementary information (SI).

3 Results and discussion

3.1 Ground State Equilibrium Geometries

A CHCl capping group breaks the symmetry of the cumulenonic chain and yields a C_s symmetry for all the Cl-cumulenone molecules analyzed in this study. In these species, the π_{CO} orbital is in-the-plane with the CHCl group for the even n and is out-of-plane for the odd n species. While $n=1-2$ have been reported in a few microwave spectroscopic studies, $n=3-4$ have not been explored. The CCSD(T)/d-aug-cc-pV(T+d)Z ground state equilibrium geometries of ClHCO and ClHC₂O compare well to the corresponding microwave spectroscopic measurements (Table 1 and Figure 1)^{47,48}. The computed and measured bond lengths agree to better than 0.01 Å for ClHCO and 0.008 Å for ClHC₂O, with $r_{C\text{Cl}}$ relatively yielding the largest discrepancies in both molecules. Theory and experiment are in better agreement for the bond angles of ClHC₂O than ClHCO with discrepancies being less than 0.14°. While a similar difference is shown for ClHCO, θ_{HCO} is an exception as the reported value is 2.80° higher than predicted at 123.490°.

Table 1 shows slight overall variations between the bond lengths and angles across the four systems, but there appear to be no particular patterns displayed in correlation with the carbon chain elongation. Although ClHC₃O and ClHC₄O geometries are not in the literature, the corresponding cumulene ketones have been reported. Notable differences are shown for θ_{OCC} and $\theta_{C_{\alpha\beta\gamma}}$ in ClHC₃O as compared to the rest of the series, where the angles are 168.992° and 138.508°, respectively. This indicates a pronounced bending at C_β, where this bend is defined in Figure 1, and another nonlinearity at C_α leading to the oxygen atom. The large deviations from linearity in the carbon chain is also reported for propadienone by previous experimental studies with $\theta_{OCC}=144.5^\circ$ and $\theta_{C_{\alpha\beta\gamma}}=169.4^\circ$ ^{73,74}.

The even n molecules, ClHC₂O and ClHC₄O, exhibit linearity in the plane (Figure 1). Unlike propadienone, the structure of butatrienone remains controversial as nonlinearity is predicted in some studies while others suggest linearity in the cumulenonic chain^{75,76}. This present study predicts a slight favor towards the C_s symmetry with a minor bend on C_β out-of-plane for ClHC₄O. This might be due to its floppy backbone with an anomalous low stretching frequency within the range of 54 - 139 cm⁻¹, depending on the levels of theory employed. This behavior will be discussed later on in the section.

In any case, the replacement of H by Cl overall yields little effect on the ground state structure of the cumulenonic chains as the odd and even n both qualitatively resemble the corresponding cu-

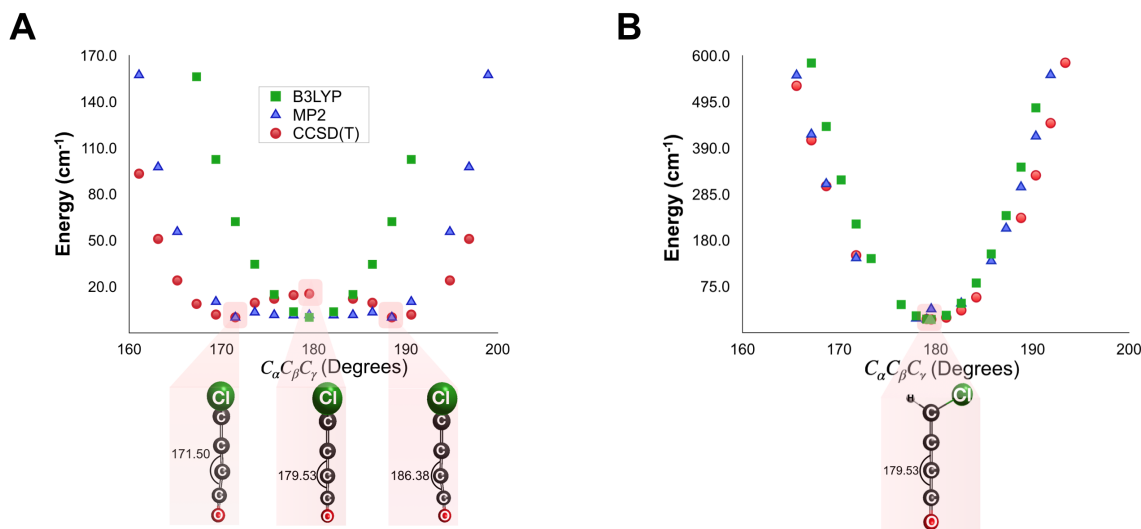


Fig. 2 Ground state potential energy curves of butatrienone chloride: (A) Out-of-plane bending mode of ClHC₄O computed at the CCSD(T), MP2, and B3LYP/d-aug-cc-pV(T+d)Z. The geometries associated with the minima and the local maximum are shown. CCSD(T) predicts a 15.357 cm^{-1} (0.001904 eV) the energy barrier from linear to bending mode. (B) In-plane bending mode of ClHC₄O calculated at the same levels of theory, where the corresponding minimum geometry is displayed. All potential energy curves are rescaled to zero minimum, where MP2 and CCSD(T) are 226,500 and 224,253 cm^{-1} above B3LYP, respectively.

mulenones. Such behavior is likely due to the minor contribution from Cl to the π -conjugation and, as such, leads to a small influence to the system. As a result, the oxygen lone-pairs dictate the nature of the backbone as seen in the case of CH₂O. Finally, the bending in both cases occur at C_β, owing to the alternating triple-single-bond resonance structures where the oxygen lone-pair donates and creates a triple bond on C_α and a lone-pair on C_β.

As compared to chloroketene, a pronounced bending on the beta carbon of propadienone chloride is predicted, indicating that the carbene character plays a more dominant role. On the other hand, propadienone has been known for its peculiar structure where an in-plane nonlinearity is observed in the allylic carbon chain⁷⁷. Despite being the most abundant species observed experimentally relative to the other two [H₂,C₃,O] isomers—cyclopropenone (*c*-C₃H₂O) and propynal (HCCCHO)—its detection is only limited to the laboratory⁷⁴. This could simply be due to a lack of interstellar data for its spectral signatures. The experimental structure of propadienone has been determined based on ground state microwave spectra of its various isotopes, where pronounced bending is reported for C_α and C_β with $\theta_{C_{\beta}C_{\alpha}O} = 169.4^\circ$ and $\theta_{C_{\alpha}C_{\beta}C_{\gamma}} = 144.5^\circ$. These measurements are confirmed with *ab initio* studies, varying from single-reference approaches such as B3LYP and the MP perturbation series to multi-reference approaches including QCISD and CASSCF^{77,78}. Similar results are also found for its radical, HCCCO, where the nonlinearity is experimentally observed for C_α and C_β, with angles 163° and 136.5°, respectively⁴⁴. As mentioned above, the structure of ClHC₃O qualitatively resembles propadienone and its radical, where the molecule is nonlinear at C_α and C_β with $\theta_{C_{\beta}C_{\alpha}O} = 169.0^\circ$ and $\theta_{C_{\gamma}C_{\beta}C_{\alpha}} = 138.5^\circ$. Additionally, θ_{HCCI} of ClHC₃O is 11° less than θ_{HCH} of propadienone, which results from the high electron density surrounding Cl that hinders it from the oxy-

gen atom as well as the cumulenic chain. The peculiar structure of ClHC₃O arises likely due to the delocalization of the in-plane π conjugation on C_β as has been observed for propadienone⁷³. The electronic structure maintains the number of in-plane bonding orbitals between ClHC₂O and ClHC₃O, but the relocation of the oxygen atom destabilizes the longer chain structure.

While propadienone possesses an in-plane nonlinearity, butatrienone is known to exist in competing symmetries of C_{2v} and C_s with both in-plane and out-of-plane bending at C_β. However, studies by microwave spectroscopy and IR Argon matrix analysis conclude that C_{2v} is the dominant form⁷⁵. The barrier between the linear and nonlinear geometries is approximately 0.0184 eV (148.4 cm^{-1}) theoretically and 0.00521 eV (42 cm^{-1}) experimentally^{76,80}. The energy barrier from C_s and C_{2v} is evidently less than the Boltzmann factor at room temperature, 0.024 eV, allowing both configurations to be populated at ambient temperature. MP3/6-31G** suggests that this molecule has a flat potential along the carbon backbone for the out-of-plane bent geometry that results from the molecule undergoing large amplitude motions⁸⁰. A recent study also shows that MP2/cc-pVTZ favors the C_s, while the DFT methods and CCSD/cc-pVTZ prefer the C_{2v} symmetry⁸¹.

Here, the CCSD(T)/d-aug-cc-pV(T+d)Z linear ground state equilibrium geometry for ClHC₄O yields an imaginary harmonic vibrational frequency corresponding to an out-of-plane bending mode at 77.6 cm^{-1} . B3LYP and MP2 computations, using the same basis set, yield a linear geometry for the lowest energy configuration that are confirmed with harmonic vibrational analyses with all real frequencies. The same out-of-plane bending frequency is computed by B3LYP and MP2 to be 65.1 cm^{-1} and 138.9 cm^{-1} , respectively. Such low stretching frequencies are indicative of the soft backbone in Cl-butatrienone that potentially allows it to interchange between nonlinearity and linearity.

Table 2 Vertical and (adiabatic) ionization energies (eV) of cumulenone chlorides with the corresponding newly creating singly-occupied orbital assignment computed at EOM-IP-CCSD level with aug-cc-pV(T+d)Z basis set in comparison with experiment^{22,79}. Vertical ionization energy calculated with d-aug-cc-pV(T+d)Z basis set are shown in the SI.

ClHC=O			ClHC=C=O			ClHC=C=C=O			ClHC=C=C=C=O		
Sym	Theor	Expt	Sym	Theor	Expt	Sym	Theor	Sym	Theor		
13a'	11.673 (11.498)	11.61±0.01 (11.51±0.01)	4a''	9.477 (8.938)	9.30±0.05 (9.11±0.03)	19a'	9.735	5a''	8.735		
3a''	12.505	12.38±0.03	15a'	12.261	12.20±0.03	18a'	10.782	20a'	11.536		
12a'	12.531	12.46±0.03	3a''	13.244	13.19±0.03	4a''	12.229	4a''	11.885		
2a''	15.712	15.28±0.01	14a'	14.854	14.70±0.05	17a'	13.639	19a'	12.008		

In an attempt to explain this behavior, electronic ground state potential energy surfaces (PES) are computed with the d-aug-cc-pV(T+d)Z basis set at the CCSD(T), MP2, and B3LYP levels of theory. The energies are calculated as a function of the $\theta_{C_{\alpha\beta\gamma}}$ using the CCSD(T) optimized geometry, where the bond angle is varied out-of-plane around its linearity at 180° with steps of 1.5° (Figure 2). As observed in regular butatrienone, the energy barriers between the two configurations calculated at CCSD(T) and MP2 are relatively small, 0.001904 eV (15.65 cm⁻¹) and 0.000192 eV (15.46 cm⁻¹), respectively. B3LYP, however, misses the barrier and predicts linearity to be the optimum configuration. The left panel of Figure 2 shows that ClHC₄O undergoes large amplitude motions, which result in a double well or a flat potential energy surface as predicted in butatrienone. Regardless, the marginal energy differences indicate that ClHC₄O essentially interchanges between the linear and the out-of-plane C_s configurations even at most interstellar temperatures above 20 K.

3.2 Ionization Energies

The vertical ionization energies (VIE) computed with the two basis sets, aug-cc-pVTZ and d-aug-cc-pVTZ, agree to 0.003 eV, indicating the diffuseness of the basis set beyond such a level does not influence the VIEs for the systems since these are pure valence structures (Table 2). In ClHCO, the lowest IE corresponds to the removal of an electron from the oxygen lone pair (n_O), followed by the chlorine lone pairs from out-of-plane (n_{Cl} (a'')) and in-plane (n_{Cl} (a')), and the π orbital (n_{π}) (Table 2). The two molecular planes are denoted as a' and a'' for in-plane and out-of-plane, respectively.

For formyl chloride, the lowest VIE, 11.673 eV, is 0.175 eV higher than the AIE optimized at the same level of theory, 11.498 eV. Both calculated values are in good agreement with the corresponding experimental measurements with 0.06 and 0.01 eV discrepancies for vertical and adiabatic transitions, respectively²². The IE_{n_O} of ClHCO is slightly higher than formic acid (HCOOH) and formaldehyde (HCOH), 11.33 and 10.88 eV, respectively⁸². The change in IE with respect to functional group substitution is therefore on the order of Cl > OH > H. The exhibited pattern is due to the anti-correlation between the amount of energy required for electron removal and the stability of the cation formed post-ionization with respect to the neutral. Evidently, the replacement of H by the more electronegative Cl destabilizes the cation and increases the IE.

A larger difference, 0.5 eV, between the lowest VIE and AIE is predicted for ClHC₂O. Table 2 shows that the calculated VIEs are

comparable with experimental measurements with differences varying between 0.05 to 0.17 eV⁷⁹. The IE_{n_O} drops by about 2.3 eV in ClHC₂O as compared to ClHCO, which is likely due to an elongation of the carbon-chain backbone enabling delocalization of the electron hole along the extended carbon chain, from O to C _{β} . The extension in the cumulenic chain also results in the second ionization to originate from the Cl lone-pair in the a' plane instead of a'' as in ClHCO.

With the exception of the first ionization, all of the successive ionizations for ClHC₃O cost less energy than those computed for ClHC₂O. On the other hand, the cation of ClHC₄O is further stabilized with respect to the parent neutral. The first ionization potential occurs at 8.735 eV, which is about 1 eV less than that of ClHC₃O. The elongation of the carbon chain again lowers the energy of the cation by enabling delocalization of the electron hole post ionization for even n molecules.

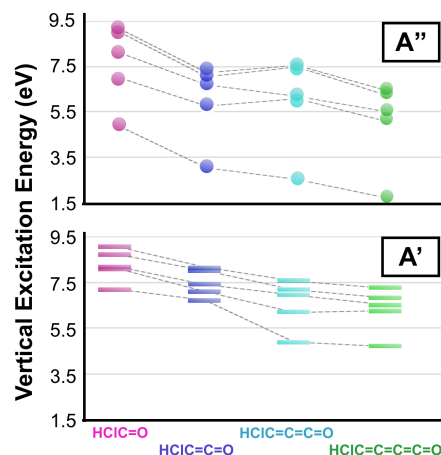


Fig. 3 Excited state properties of cumulenone chlorides: Vertical excitation energies with the corresponding characters of formyl chloride, chloroketene, propadienone chloride, and butatrieneone computed at the EOM-CCSD/d-aug-cc-pV(T+d)Z level for the (top) out-of-plane and the (bottom) in-plane transitions.

3.3 Vertical Excitation Energies

Small chlorinated molecules, such as chloromethane and hydrogen chloride, are used to benchmark the adequacy of the Cl atom treatment in EOM-CCSD with different basis sets. The results are shown in Tables 13 & 14 in the SI. The excited state energies agree with the experimental measurements to within 0.1 - 0.15 eV of experiment^{83,84}, which supports the claim that this

method is an appropriate choice. Moreover, the T_1 diagnostic confirms that EOM-CCSD provides adequate predictions for the excited states and multi-reference dynamical treatments are not necessary. T_1 diagnostics are less than 0.02 for all species and particularly 0.016 for formyl chloride, which satisfies the typical multi-reference character, 0.02, limit⁸⁵. The $+spd$ technique^{68,70} also produces results that are comparable to those of the d- and t-aug-cc-pV(T+d)Z for the CCSD(T), B3LYP, and MP2 optimized geometries (refer to Tables 19, 22, 25 in the SI) indicating that its usage is also proper when necessary.

Table 3 Excited state energies of cumulenone chlorides computed at EOM-CCSD level using ground state equilibrium geometries optimized at CCSD(T)/d-aug-cc-pV(T+d)Z with Dunning’s basis sets where SA is the state assignment, E is the vertical excitation energy in eV and f is the oscillator strength. The last two columns report the energies and oscillator strengths computed with t-aug-cc-pV(T+d)Z for ClHCO and ClHC₂O and aug-cc-pVTZ+ spd for ClHC₃O and ClHC₄O.

SA	E	f	E	f	E	f
	aug-cc-pV(T+d)Z		d-aug-cc-pV(T+d)Z		t-aug-cc-pV(T+d)Z	
HCIC=O						
1 $^1A''$	4.963	0.1824	4.951	0.0001	4.951	0.0001
2 $^1A''$	7.091	0.2606	7.069	0.0003	7.069	0.0003
1 $^1A'$	7.245	0.2663	7.216	0.0191	7.216	0.0191
2 $^1A'$	8.166	0.3001	8.143	0.0617	8.143	0.0618
HCIC=C=O						
1 $^1A''$	3.057	0.0000	3.057	0.0000	3.057	0.0000
2 $^1A''$	5.759	0.0043	5.749	0.0040	5.749	0.0040
3 $^1A''$	6.753	0.0015	6.692	0.0020	6.692	0.0020
1 $^1A'$	6.730	0.0019	6.729	0.0018	6.729	0.0018
	aug-cc-pV(T+d)Z		d-aug-cc-pV(T+d)Z		aug-cc-pVTZ+ spd	
HCIC=C=C=O						
1 $^1A''$	2.560	0.0002	2.559	0.0002	2.559	0.0002
1 $^1A'$	4.903	0.0813	4.902	0.0809	4.902	0.0812
2 $^1A''$	6.103	0.0000	6.103	0.0000	6.103	0.0000
3 $^1A''$	6.208	0.0023	6.205	0.0023	6.207	0.0023
HCIC=C=C=C=O						
1 $^1A''$	1.731	0.0000	1.731	0.0000	1.731	0.0000
1 $^1A'$	4.723	0.0431	4.722	0.0446	4.722	0.0431
2 $^1A''$	5.085	0.0001	5.085	0.0002	5.085	0.0001
3 $^1A''$	5.544	0.0016	5.516	0.0016	5.520	0.0016

Furthermore, the method employed to compute the reference geometry is largely irrelevant. Vertical excitation energies computed at the EOM-CCSD level with different basis sets for the different optimized geometries are displayed in Figure 4 in the SI for each of the molecules in question. The corresponding energies are reported in Tables 15-27 in the SI. The comparison of the EOM-CCSD/d-aug-cc-pV(T+d)Z excited state energies computed for the three different optimized geometries for each molecule is shown. The discrepancies are largely insignificant being smaller than 0.1 eV for ClHCO, ClHC₂O, and ClHC₄O, and they are around 0.1 eV for ClHC₃O. The 3 $^1A''$ state in ClHC₃O is an exception, which displays apparent differences between the three methods with 6.205, 6.465, 6.371 eV for CCSD(T), B3LYP, and MP2, respectively. The largest difference among different methods is 0.260 eV between B3LYP and CCSD(T) for ClHC₃O, but this is an outlier. Hence, the choice of method for computing the reference geometry does not have a notable effect on the excitation energy into the VUV.

The EOM-CCSD vertical excitation energy decreases with re-

spect to the extension of the cumulenonic chain for the transitions originating from the plane as shown in Figure 3. The A'' transitions in the cumulenone chloride series also adapt the same pattern with an exception of ClHC₃O. Excluding the lowest A'' excitation, promotion of an electron out of the plane requires more energy for ClHC₃O than ClHC₂O.

3.3.1 Formyl Chloride

Chloromethane (CH₃Cl) and formaldehyde (H₂CO) are well investigated in both experiment and theory. Investigations on the Rydberg series of these molecules provide that CH₃Cl and C=O functional groups have their Rydberg series start around 8.81 and 7.09 eV, respectively^{84,86}. By containing both of these groups, the Rydberg series of ClCHO is expected to occur in the vicinity of 8.0 eV. While the Ahlrichs’s and correlation consistent sets are largely self-consistent, the discrepancies between corresponding transition energy between basis classes are significant with differences from 0.02 eV to 1.61 eV for higher-energy transitions. Since EOM-CCSD is a variational method and the correlation consistent basis provide much lower values for the energy levels, the correlation consistent energies are concluded to be more reliable. Hence, the correlation consistent basis sets are used to predict the properties of the rest of the Cl-cumulenone series.

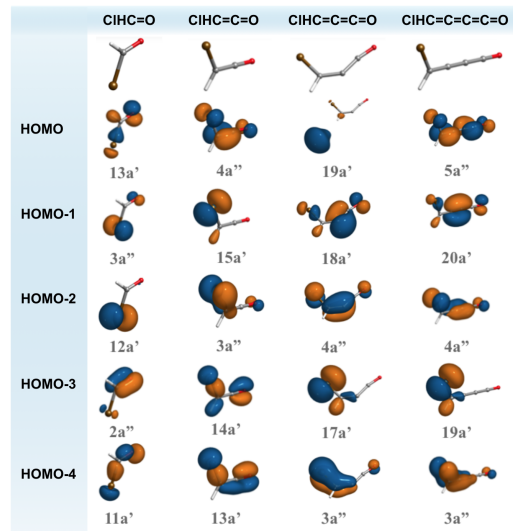


Fig. 4 Molecular orbitals of cumulenone chlorides: Highest occupied molecular orbitals (HOMO) of ClHCO, ClHC₂O, ClHC₃O, ClHC₄O computed at HF-SCF/t-aug-cc-pV(T+d)Z level of theory.

In ClCHO, there is a non-negligible (0.085 eV) decrease in the excitation energy computed with the aug-cc-pV(T+d)Z basis set as compared to that from d-aug-cc-pV(T+d)Z for the 4 $^1A'$ state and higher. The energies converge when more diffuse functions are added with the t-aug-cc-pV(T+d)Z basis as they match the doubly augmented basis to the significant digits reported here implying Rydberg-character in this state. This trend continues for the higher-energy states of ClCHO. As a result, the following discussion will be based upon EOM-CCSD/t-aug-cc-pV(T+d)Z results.

Consistent with the order of ionizations, the highest occupied molecular orbitals (HOMO) is the oxygen lone pair, n_O (13a’);

(HOMO-1) is an out-of-plane chlorine lone pair, n_{Cl} ($3a''$); and (HOMO-2) is an in-plane chlorine lone pair n_{Cl} ($12a'$) (Figure 4). The first transition at 4.951 eV is an $\pi^*_{CO} \leftarrow n_O$ transition, which is optically dark, where π^* orbital is localized on the carbonyl with minor diffuse character on Cl (Figure 5 & Table 3). EOM-CCSD predicts the transition into the first excited state, $1^1A''$, to be 4.951 eV, which is 0.089 eV lower than the result computed with MRD-CI+Q/cc-pVDZ+*spd*, 5.04 eV⁸⁷. The EOM-CCSD energy, however, is closer to the gas phase experimental measurement at 4.86 eV⁸⁸. The discrepancies between EOM-CCSD and MRD-CI+Q/cc-pVDZ+*spd* vary from 0.083 - 0.916 eV, where the largest difference is for the $3^1A''$ state where, again, the present EOM-CCSD results are taken to be more reliable from the previous and current (above) benchmarks (Table 15 in the SI).

The next two lowest excited states, $1^1A'$ and $2^1A''$, share the same electronic excitation character as $1^1A''$ (Figure 5), indicating that these three come as a set of excitations out of degenerate p orbital groupings from the chlorine atom. However, $1^1A'$ is marginally brighter than the other two with $f = 0.0191$ due to the population of electrons being excited not only from the n_{Cl} (a'') but also from the n_{Cl} (a'). This increases the orbital overlap as well as the brightness of the transition. Meanwhile, the $1^1A''$ and $2^1A''$ states have electrons excited out of n_O (a') and n_{Cl} (a') orbitals to out-of-plane orbitals. The poor overlap between the occupied and the virtual molecular orbitals renders these transitions optically dark.

The next set of transitions, $3^1A'$, $4^1A'$, and $3^1A''$, share the same σ^* electronic character with diffuse characters located on Cl (Figure 5). The $4^1A'$ state is the only bright transition with $f = 0.3407$, owing to the same symmetry excitation from n_O ($13a'$) to σ^*_{CCl} . These states are much closer in energy than the previous set where the order of the orbitals out of which the excitation occurs is reversed. Specifically, the lowest excitation originates from n_{Cl} ($3a''$), and the highest transition involves n_{Cl} ($13a'$).

The Rydberg series of ClCHO starts at 8.143 eV close to the 8.0 eV point hypothesized above from averaging the Rydberg excitations of Cl-bearing and C=O-containing species. This series is also the third trio of states and consists of the $2^1A'$, $4^1A''$, and $5^1A'$ states. The $2^1A'$ and $4^1A''$ states lie 1 eV apart, where diffuse orbitals are localized around the HC-Cl bond angle from Figure 5. The $4^1A''$ and $5^1A'$ states are nearly degenerate at 9.099 and 9.097 eV with comparable oscillator strengths of 0.0532 and 0.0485, respectively. Both of these states are relatively dim as compared to the $4^1A'$. The $2^1A'$ state is marginally brighter due to the same symmetry excitation and better orbital overlap of the occupied and virtual orbitals. Although they appear to be different, $5^1A'$ and $4^1A''$ are topologically similar as they share similar valence character again shown in Figure 5. Thus the, $5^1A'$ fits in well with this set owing to its electronic character. It is, however, much more diffuse with an s -Rydberg present in the transition. The diffuseness leads to a small discrepancy in the energies of the two states since diffuse electron density tends to lower the energy of the state⁶⁶.

Since there is some valence character present in addition to diffuse characters, this trio of states as well as the $5^1A''$ state provide for the lowest energy valence-Rydberg mixing transitions

as exhibited in Figure 5. States $5^1A'$ and $5^1A''$ have the most prominent diffuse character of orbitals, where the $5^1A'$ state exhibits s -Rydberg mixing at 9.097 eV and the $5^1A''$ state shows p -Rydberg character. The Rydberg series in ClCHO is approximately 1.0 eV higher than formaldehyde, which occurs at 7.08 eV⁸⁶.

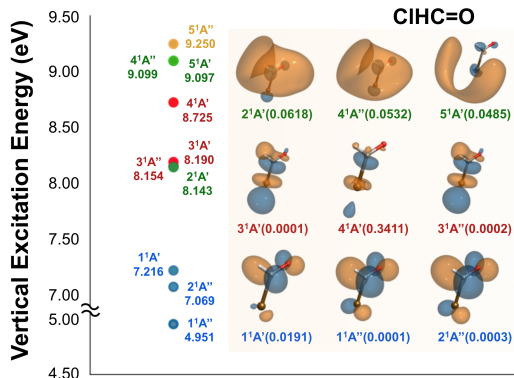


Fig. 5 Excited state properties of formyl chloride: The vertical excitation energies with the corresponding symmetry computed at EOM-CCSD/t-aug-cc-pV(T+d)Z are shown on the left. The color code—blue, red, green—distinguishes different sets of transitions. The excited state transition orbitals of each set are displayed on the right panel with the corresponding color code. The molecular orbital isosurfaces are 0.03, 0.02, 0.01 e^2/Bohr for the bottom, middle, and top group of transitions.

3.3.2 Chloroketene

The aug-cc-pV(T+d)Z energies are consistently 0.1 eV larger than the energies computed with the d-aug-cc-pV(T+d)Z basis up to $6^1A''$ (Table 3). More details can be found in Table 19 in the SI. The energy gap between the two bases rises to 0.5 - 0.6 eV starting at $7^1A''$ (~ 7.9 eV) due to the significant increase in the diffuseness of the electrons, an indication of Rydberg character. This pattern is only predicted for states that are higher than 8 eV even though the Rydberg series starts at 5.749 eV for $2^1A''$. This pattern is not quite as pronounced in the case of ClCHO where the largest shift is approximately 0.2 - 0.3 eV. However, and like ClCHO, the t-aug-cc-pV(T+d)Z basis set is used for discussion here since the properties are considered converged for such diffuseness.

The first transition of ClHC₂O at 3.057 eV is 1.894 eV lower than that of ClCHO. $1^1A'$, $1^1A''$, and $6^1A''$ form a trio of states that, again, possess the same qualitative topology with electrons being excited out of the π ($4a''$), n_{Cl} ($15a'$), and $n_{Cl,O}$ ($3a''$) orbitals (Figure 4). While $1^1A''$ is completely dark, $1^1A'$ has a marginally larger oscillator strength than the $6^1A''$ state due to better overlap from occupied and virtual orbitals of the same symmetry. The energy of the $6^1A''$ is lowered by 2.3 eV in ClHC₂O as compared to the corresponding transition in ClCHO, allowing it to form a set with $1^1A'$ and $1^1A''$ instead of $2^1A''$ in the case of ClCHO. While the $2^1A''$ excitation only originates from π ($4a''$), the transition into the $6^1A''$ state involves participation from both the π ($4a''$) and $n_{Cl,O}$ ($3a''$) orbitals (Figure 6).

Besides $3^1A'$, the rest of the chloroketene excited state transitions are diffuse Rydberg mixing with valence character. The

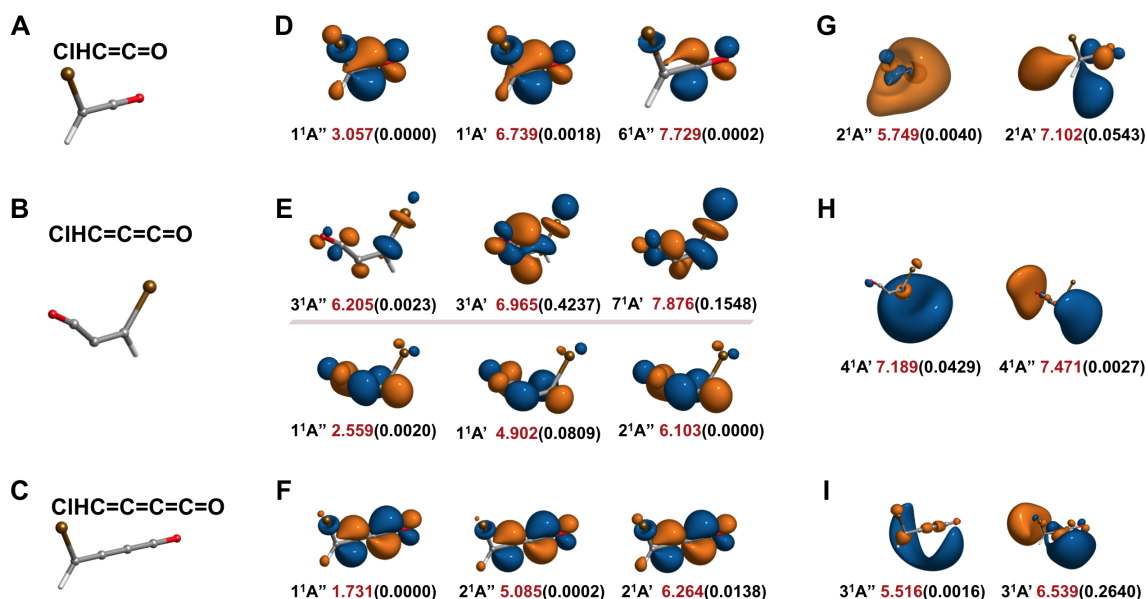


Fig. 6 Excited state transition orbitals: Ground state geometries of (A) chloroketene, (B) propadienone chloride, (C) butatrienone chloride. The group(s) of orbitals with the same character for each molecule computed at CCSD(T)/t-aug-cc-pV(T+d)Z are shown in panels (D), (E), (F), respectively. The lowest Rydberg–valence transitions of two symmetries for each molecule are shown in panels (G), (H), (I). The state assignment, associated vertical excitation energy in red, and oscillator strength in parentheses of each transition are specified underneath. The isosurfaces (iso) used for plotting the excited state molecular orbitals for ClHC₂O are (D) 0.035 and (G) 0.005 (left) and 0.011 (right) e²/Bohr, for ClHC₃O are (E) 0.025, 0.0125, (H) 0.040 e²/Bohr, and for ClHC₄O are (F) 0.035 and (I) 0.0125 e²/Bohr.

Rydberg series starts with the 2 ¹A'' transition with electrons being excited out of the π (4a'') cloud, resulting in valence orbitals mixing with an *s*–Rydberg localized on the C=O bond (Figure 6G). Meanwhile, 2 ¹A' at 7.102 eV is the lowest Rydberg–valence state—with a *p*–Rydberg localized on C _{β} —for the in-plane transitions. Rydberg character in chloroketene starts at 2 eV lower than in formyl chloride, which appears at around 8 and 9 eV for the A' and A'' series, respectively (Figure 6G). The dramatic increase in the diffuse character in ClHC₂O is exhibited through the dropping of vertical excitation energies with more diffuse basis sets and the orbital characters. Overall, the computed properties of the lowest 15 excited states yield no evidence of the Rydberg series converging to the excited states of the cation.

3.3.3 Propadienone Chloride

Propadienone is known for its peculiar structure where an in-plane nonlinearity is observed in the allylic carbon chain⁷³. The addition of two carbons along with chlorine substitution raises the cost of the calculations relative to the benchmarks from formyl chloride. Therefore, the TZ+*spd* basis sets are employed here where computational cost is reduced with little-to-no sacrifice in accuracy^{68,70}. This basis set is benchmarked for ClHCO in the SI. There are no dramatic differences in the TZ+*spd* basis compared with t-aug-cc-pV(T+d)Z. Furthermore, shifts from the aug-cc-pVTZ excited state energy compared with either TZ+*spd* or t-aug-cc-pV(T+d)Z are consistent. Hence, this basis is utilized for the most diffuse computations in ClHC₃O.

Figures 5 and 6E show that lowest two groups of excitations in ClHCO and ClHC₃O share similar electronic characters. Specifically, the 1 ¹A'', 1 ¹A', and 2 ¹A'' states form a set in both ClHCO

and ClHC₃O where 1 ¹A' is the brightest due to a better orbital overlap of the same symmetry excitation. The π -cloud localized along the carbon chain is also extended to the oxygen atom with minor diffuse character on the chlorine atom. This same behavior is present in chloroketene where the π cloud is extended. The electron densities in these excitations originate from HOMO = *n*_O (18a'), (HOMO-1) = *n*_{Cl} (4a''), and (HOMO-2) = *n*_{Cl} (17a') (Figure 4). The 1 ¹A' state results from a mixed transition with electron densities being excited from mostly *n*_{Cl} (4a'') with a minor contribution from *n*_O (18a'). The greater orbital overlaps render it a dim transition as compared to the other two. Moreover, the electronic character of state 8 ¹A' also seems to resemble this group with an addition of some *s*–Rydberg (Figure 2 & Table 22 in SI).

The second trio of near-degenerate states consists of the 3 ¹A', 3 ¹A'', and 7 ¹A', where 3 ¹A'' is the darkest state due to the poor orbital overlap resulting from out-of-plane excitations. A non-bonding orbital along the CCl bond is seen for both ClHCO and ClHC₃O. Moreover, state 3 ¹A' is also the brightest among all the excitations reported for this molecule with *f* = 0.4237. This excitation is contaminated by some population from the (HOMO-1) in addition to the HOMO. Meanwhile, 7 ¹A' involves a minor contribution from the (HOMO-3) in addition to the (HOMO-2) and is also a bright transition.

The lowest Rydberg–valence transitions in both symmetries appear for the 4 ¹A' (7.189 eV) and 4 ¹A'' (7.471 eV) states (Figure 6H). In both cases, the *s*–Rydberg and *p*–Rydberg are dominant, respectively. The energy gap of these two Rydberg states is 0.282 eV, which is slightly larger than the energy difference of the lowest Rydberg states of the two symmetries in ClCHO, 0.011 eV.

Both of these differences are much smaller than that of ClHC₂O, which is 1.302 eV.

3.3.4 Butatrienone Chloride

Even though butatrienone interchanges between the linear and out-of-plane *C_s* configurations, such relatively small differences in the structure of the molecule do not change its orbital topology by any meaningful amount. Hence, the actual structure of the backbone will be insignificant for computing meaningful properties, such as electronic excitation energies, ionization energies, and oscillator strengths. In point, the excited state energies computed at the EOM-CCSD for three optimized geometries, even with some bent and other linear, are in excellent agreement as shown in Figure 8 in the SI. Consequently, the linear configuration will be used for all vertical excitation energy computations.

There is a similarity between the first set of excitations in the even *n* species (Figure 6D & E). The excitation characters of the first set in ClHC₄O closely resemble those of ClHC₂O where 1 ¹A' and 6 ¹A'' are replaced by 2 ¹A' and 2 ¹A'', respectively. The elongation of the conjugated system stabilizes the entire molecule and, thus, significantly lowers the energy for the last component of this set.

The lowest Rydberg–valence transitions in both symmetries are found for the 3 ¹A' (6.539 eV) and 3 ¹A'' (5.516 eV) states (Figure 6I). The 3 ¹A'' state is dominated with *s*–Rydberg character while the 3 ¹A' state is surrounded by a *p*–Rydberg orbital with π valence character. The energy gap between the two lowest Rydberg–valence transitions is 1.023 eV, which is 0.3 eV lower than that of ClHC₂O. The transition into the 3 ¹A' state has a significantly brighter oscillator strength, 0.2640, due to a much greater orbital overlap. The Rydberg–valence states of the two symmetries of the even *n* species seem to have a bigger energy difference (more than 1.0 eV) as compared to the odd *n* molecules which is less than 0.3 eV. The extension of the carbon chain not only decreases the energy levels for most states but also lowers the Rydberg series. Similar to ClHC₂O, there is approximately a 0.5 eV difference in energy starting with the transition into the 5 ¹A' state, where significant diffuse character starts to appear (Figure 3 and Table 25 in the SI).

4 Conclusions

The behavior of the cumulenone chlorides is in many ways an average of the effects observed in regular cumulenone molecules and organochlorides. The structures of formyl chloride and chloroketene strongly resemble the corresponding hydrocarbon species, indicating that the chlorine atom has a minor contribution towards the chemical interactions that determine the bonding structure of the system. Meanwhile, the chlorine replacement perturbs the structures of ClHC₃O and ClHC₄O and leads to slight differences to a much greater extent as compared to the corresponding cumulenone species. The substitution effect is displayed via the discrepancies in the ionization energies for the chlorinated series as they are approximately 1 eV higher than the cumulenones. This is likely caused by the cation—formed post-ionization—becoming less stable due to the presence of the chlorine atom. The carbene character is enhanced in ClHC₃O as com-

pared to H₂CHO as more bending is observed on the beta carbon. Differently, ClHC₄O, with a soft backbone, undergoes low-energy large amplitude motion and interchanges between out-of-plane bending and linearity. Even so, the geometries and ionization energies computed here for formyl chloride and chloroketene are in excellent agreement with microwave spectroscopic measurements indicating that such computed properties are likely valid for these as-of-yet unobserved species.

Similar to the ionization energies, the Rydberg series of the chlorinated systems also starts about 1 eV higher than the cumulenone family. Electronic excitations from the two chlorine lone-pairs and an oxygen lone-pair form different groups of excitations that consist of three different states sharing the same character originating from the valence *p* orbitals on the chlorine atom. These transitions may ultimately lead to a chain of reactions that are directly relevant to atmospheric science and astrochemistry. Halochemistry of the former is well-established, but that of the latter is just emerging. The kinetics for the photolysis and the successive collisions between the resulting radicals and other species can be further investigated experimentally as has been done recently¹⁶. Hence, the obtained photophysical properties of these systems will potentially enable the modeling of such reactions that occur in both the stratosphere and the ISM.

Conflicts of interest

The authors have no conflicts to declare.

Acknowledgements

Q.L.N and W.K.P acknowledges support from Department of Energy Office of Basic Energy Sciences (DE-SC0012628). Q.L.N. was also supported by the National Science Foundation Graduate Research Fellowship (DGE-1144083). R.C.F. acknowledges support from the National Aeronautics and Space Administration (NNX17AH15G) and from the National Science Foundation (OIA-1757220). We thank Michael Tanksavala for helpful discussions and Prof. Margaret M. Murnane and Prof. G. Barney Ellison for the support of this work and comments on the manuscript.

Notes and references

- 1 J. Berkowitz, *Photoabsorption, Photoionization, and Photoelectron Spectroscopy*, Academic Press, Newyork, 1979.
- 2 R. W. Shaw, T. B. Brill and D. L. Thompson, *Overviews of recent research on energetic materials*, World Scientific, Singapore, 2005.
- 3 D. D. Dlott, *Energetic Materials*, Elsevier, 2003, vol. 13, pp. 125 – 191.
- 4 M. Kuklja, *Applied Physics A*, 2003, **76**, 359–366.
- 5 D. R. Marsh, in *Chemical–Dynamical Coupling in the Mesosphere and Lower Thermosphere*, ed. M. A. Abdu and D. Pancheva, Springer Netherlands, Dordrecht, 2011, pp. 3–17.
- 6 H. Lühr, H. Liu, J. Park and S. Müller, in *New Aspects of the Coupling Between Thermosphere and Ionosphere, with Special regards to CHAMP Mission Results*, ed. M. A. Abdu and

- D. Pancheva, Springer Netherlands, Dordrecht, 2011, pp. 303–316.
- 7 C. He, S. M. Hörst, N. K. Lewis, J. I. Moses, E. M.-R. Kempton, M. S. Marley, C. V. Morley, J. A. Valenti and V. Vuitton, *ACS Earth and Space Chemistry*, 2019, **3**, 39–50.
 - 8 K. I. Öberg, *Chemical Reviews*, 2016, **116**, 9631–9663.
 - 9 D. K. Böhme, *Phys. Chem. Chem. Phys.*, 2011, **13**, 18253–18263.
 - 10 H. Reisler and A. I. Krylov, *International Reviews in Physical Chemistry*, 2009, **28**, 267–308.
 - 11 M. B. Robin, *Higher Excited States of Polyatomic Molecules*, Academic Press, 1970.
 - 12 T. Mori, W. J. Glover, M. S. Schuurman and T. J. Martinez, *The Journal of Physical Chemistry A*, 2012, **116**, 2808–2818.
 - 13 E. G. Champenois, N. H. Shivaram, T. W. Wright, C.-S. Yang, A. Belkacem and J. P. Cryan, *The Journal of Chemical Physics*, 2016, **144**, 014303.
 - 14 T. Geng, O. Schalk, S. P. Neville, T. Hansson and R. D. Thomas, *The Journal of Chemical Physics*, 2017, **146**, 144307.
 - 15 S. P. Neville and G. A. Worth, *The Journal of Chemical Physics*, 2014, **140**, 034317.
 - 16 W. K. Peters, D. E. Couch, B. Mignolet, X. Shi, Q. L. Nguyen, R. C. Fortenberry, H. B. Schlegel, F. Remacle, H. C. Kapteyn, M. M. Murnane and W. Li, *Proceedings of the National Academy of Sciences*, 2017, E11072–E11081.
 - 17 D. W. Turner, *Molecular Photoelectron Spectroscopy: A Handbook of He 584 Å Spectra*, Wiley Interscience, London, 1970.
 - 18 B. R. Lewis, A. N. Heays, S. T. Gibson, H. Lefebvre-Brion and R. Lefebvre, *The Journal of Chemical Physics*, 2008, **129**, 164306.
 - 19 H. Lefebvre-Brion and M. Eidelsberg, *Journal of Molecular Spectroscopy*, 2012, **271**, 59 – 65.
 - 20 M. O. Vieitez, T. I. Ivanov, C. A. de Lange, W. Ubachs, A. N. Heays, B. R. Lewis and G. Stark, *The Journal of Chemical Physics*, 2008, **128**, 134313.
 - 21 G. Schönnenbeck, H. Biehl, F. Stuhl, U. Meier and V. Staemmler, *The Journal of Chemical Physics*, 1998, **109**, 2210–2219.
 - 22 D. Frost, C. McDowell and N. Westwood, *Chemical Physics Letters*, 1977, **51**, 607 – 610.
 - 23 Q. S. Li and Q. Luo, *The Journal of Physical Chemistry A*, 2003, **107**, 10435–10440.
 - 24 G. Fantechi, N. R. Jensen, O. Saastad, J. Hjorth and J. Peeters, *Journal of Atmospheric Chemistry*, 1998, **31**, 247–267.
 - 25 V. Catoire, R. Lesclaux, W. F. Schneider and T. J. Wallington, *The Journal of Physical Chemistry*, 1996, **100**, 14356–14371.
 - 26 T. J. Wallington, M. D. Hurley and W. F. Schneider, *Chemical Physics Letters*, 1996, **251**, 164 – 173.
 - 27 B. W. Gay, P. L. Hanst, J. J. Bufalini and R. C. Noonan, *Environmental Science & Technology*, 1976, **10**, 58–67.
 - 28 J. P. Grotzinger, J. Crisp, A. R. Vasavada, R. C. Anderson, C. J. Baker, R. Barry, D. F. Blake, K. S. Conrad, Pamela aand Edgett, B. Ferdowski, R. Gellert, J. B. Gilbert, M. Golombek, J. Gómez-Elvira, D. M. Hassler, L. Jandura, M. Litvak, P. Mahaffy, J. Maki, M. Meyer, M. C. Malin, I. Mitrofanov, J. J. Simmonds, D. Vaniman, R. V. Welch and R. C. Wiens, *Space Science Reviews*, 2012, **170**, 1572–9672.
 - 29 K. Biemann, J. Oro, P. Toulmin, L. E. Orgel, A. O. Nier, D. M. Anderson, P. G. Simmonds, D. Flory, A. V. Diaz, D. R. Rushneck and J. A. Biller, *Science*, 1976, **194**, 72–76.
 - 30 Lis, D. C., Pearson, J. C., Neufeld, D. A., Schilke, P., Müller, H. S. P., Gupta, H., Bell, T. A., Comito, C., Phillips, T. G., Bergin, E. A., Ceccarelli, C., Goldsmith, P. F., Blake, G. A., Bacmann, A., Baudry, A., Benedettini, M., Benz, A., Black, J., Boogert, A., Bottinelli, S., Cabrit, S., Caselli, P., Castets, A., Caux, E., Cernicharo, J., Codella, C., Coutens, A., Crimier, N., Crockett, N. R., Daniel, F., Demyk, K., Dominic, C., Dubernet, M.-L., Emprechtinger, M., Encrenaz, P., Falgarone, E., Fuente, A., Gerin, M., Giesen, T. F., Goicoechea, J. R., Helmich, F., Hennebelle, P., Henning, Th., Herbst, E., Hily-Blant, P., Hjalmarson, Å., Hollenbach, D., Jack, T., Joblin, C., Johnstone, D., Kahane, C., Kama, M., Kaufman, M., Klotz, A., Langer, W. D., Larsson, B., Le Boulrot, J., Lefloch, B., Le Petit, F., Li, D., Liseau, R., Lord, S. D., Lorenzani, A., Maret, S., Martin, P. G., Melnick, G. J., Menten, K. M., Morris, P., Murphy, J. A., Nagy, Z., Nisini, B., Ossenkopf, V., Pacheco, S., Pagani, L., Parise, B., Pérault, M., Plume, R., Qin, S.-L., Roueff, E., Salez, M., Sandqvist, A., Saraceno, P., Schlemmer, S., Schuster, K., Snell, R., Stutzki, J., Tielens, A., Trappe, N., van der Tak, F. F. S., van der Wiel, M. H. D., van Dishoeck, E., Vastel, C., Viti, S., Wakelam, V., Walters, A., Wang, S., Wyrowski, F., Yorke, H. W., Yu, S., Zmuidzinas, J., Delorme, Y., Desbat, J.-P., Güsten, R., Krieg, J.-M. and Delforge, B., *Astronomy & Astrophysics*, 2010, **521**, L9.
 - 31 D. A. Neufeld, E. Roueff, R. L. Snell, D. Lis, A. O. Benz, S. Bruderer, J. H. Black, M. D. Luca, M. Gerin, P. F. Goldsmith, H. Gupta, N. Indriolo, J. L. Boulrot, F. L. Petit, B. Larson, G. J. Melnick, K. M. Menten, R. Monje, Z. Nagy, T. G. Phillips, A. Sandqvist, P. Sonnentrucker, F. van der Tak and M. G. Wolfire, *The Astrophysical Journal*, 2012, **748**, 37.
 - 32 M. D. Luca, H. Gupta, D. Neufeld, M. Gerin, D. Teyssier, B. J. Drouin, J. C. Pearson, D. C. Lis, R. Monje, T. G. Phillips, J. R. Goicoechea, B. Godard, E. Falgarone, A. Coutens and T. A. Bell, *The Astrophysical Journal Letters*, 2012, **751**, L37.
 - 33 W.-H. Fang and R.-Z. Liu, *The Journal of Chemical Physics*, 2001, **115**, 10431–10437.
 - 34 W. M. Irvine, L. W. Avery, P. Friberg, H. E. Matthews and L. M. Ziurys, *Astrophysical letters and communications*, 1988, **26**, 167a180.
 - 35 M. Guelin, N. N. and J. Cernicharo, *Astronomy and Astrophysics*, 1998, **335**, L1–L4.
 - 36 J. Cernicharo, A. M. Heras, A. G. G. M. Tielens, J. R. Pardo, F. Herpin, M. Guélin and L. B. F. M. Waters, *The Astrophysical Journal*, 2001, **546**, L123–L126.
 - 37 L. M. Ziurys, *Proceedings of the National Academy of Sciences*, 2006, **103**, 12274–12279.
 - 38 A. Inness, F. Baier, A. Benedetti, I. Bouarar, S. Chabrilat, H. Clark, C. Clerbaux, P. Coheur, R. J. Engelen, Q. Errera, J. Flemming, M. George, C. Granier, J. Hadji-Lazaro, V. Huijnen, D. Hurtmans, L. Jones, J. W. Kaiser, J. Kapsomenakis,

- K. Lefever, J. Leitão, M. Razinger, A. Richter, M. G. Schultz, A. J. Simmons, M. Suttie, O. Stein, J.-N. Thépaut, V. Thouret, M. Vrekoussis, C. Zerefos and the MACC team, *Atmospheric Chemistry and Physics*, 2013, **13**, 4073–4109.
- 39 N. Washida, S. Hatakeyama, H. Takagi, T. Kyogoku and S. Sato, *The Journal of Chemical Physics*, 1983, **78**, 4533–4540.
 - 40 L. E. Snyder, D. Buhl, B. Zuckerman and P. Palmer, *Physical Review Letters*, 1969, **22**, 679–681.
 - 41 W. M. Irvine, R. D. Brown, D. M. Cragg, P. Friberg, P. D. Godfrey, N. Kaifu, H. E. Matthews, M. Ohishi, H. Suzuki and H. Takeo, *The Astrophysical Journal Letters*, 1988, **335**, L89.
 - 42 H. E. Matthews, W. M. Irvine, P. Friberg, R. D. Brown and P. D. Godfrey, *Nature*, 1984, **310**, 125–126.
 - 43 K. B. Krauskopf and G. K. Rollefson, *Journal of the American Chemical Society*, 1934, **56**, 2542–2548.
 - 44 A. L. Cooksy, J. K. G. Watson, C. A. Gottlieb and P. Thaddeus, *The Journal of Chemical Physics*, 1994, **101**, 178–186.
 - 45 R. S. Timonen, E. Ratajczak and D. Gutman, *The Journal of Physical Chemistry*, 1988, **92**, 651–655.
 - 46 H. Niki, P. Maker, L. Breitenbach and C. Savage, *Chemical Physics Letters*, 1978, **57**, 596 – 599.
 - 47 M. C. L. Gerry, W. Lewis-Bevan and N. P. C. Westwood, *The Journal of Chemical Physics*, 1983, **79**, 4655–4663.
 - 48 H. Takeo and C. Matsumura, *The Journal of Chemical Physics*, 1976, **64**, 4536–4540.
 - 49 E. C. Fayolle, K. I. Öberg, J. K. Jørgensen, K. Altwegg, H. Calcutt, H. S. P. Müller, M. Rubin, M. H. D. van der Wiel, P. Bjerkeli, T. L. Bourke, A. Coutens, E. F. van Dishoeck, M. N. Drozdovskaya, R. T. Garrod, N. F. W. Ligterink, M. V. Persson and S. F. Wampfler, *Nature Astronomy*, 2017, **1**, 703–708.
 - 50 M. H. Hecht, S. P. Kounaves, R. C. Quinn, S. J. West, S. M. M. Young, D. W. Ming, D. C. Catling, B. C. Clark, W. V. Boynton, J. Hoffman, L. P. DeFlores, K. Gospodinova, J. Kapit and P. H. Smith, *Science*, 2009, **325**, 64–67.
 - 51 J. F. Stanton, J. Gauss, M. E. Harding, , P. G. Szalay, with Contributions from A. A. Auer, R. J. Bartlett, U. Benedikt, C. Berger, D. E. Bernholdt, Y. J. Bomble, L. Cheng, O. Christiansen, M. Heckert, O. Heun, C. Huber, T. Jagau, D. Jonsen, J. Juselius, K. Klein, W. J. Lauderdale, D. A. Matthews, T. Metzroth, L. A. Muck, D. P. O'Neill, D. R. Price, E. Prochnow, C. Puzzarini, K. Ruud, F. Schiffmann, W. Schwalbach, C. Simmons, S. Stopkowicz, A. Tajti, J. Vazquez, F. Wang, J. D. Watts, the Integral Packages MOLECULE (J. Almlof, P. R. Taylor), P. P. R. Taylor), A. T. Helgaker, H. J. A. Jensen, P. Jørgensen, J. Olsen), E. R. by A. V. Mitin and C. van Wüllen, *CFOUR, a Quantum Chemical Program Package. For the current version, see <http://www.cfour.de>*.
 - 52 J. M. Turney, A. C. Simmonett, R. M. Parrish, E. G. Hohenstein, F. A. Evangelista, J. T. Fermann, B. J. Mintz, L. A. Burns, J. J. Wilke, M. L. Abrams, N. J. Russ, M. L. Leininger, C. L. Janssen, E. T. Seidl, W. D. Allen, H. F. Schaefer, R. A. King, E. F. Valeev, C. D. Sherrill and T. D. Crawford, *Wiley Interdisciplinary Reviews: Computational Molecular Science*, 2012, **2**, 556–565.
 - 53 T. H. Dunning, *The Journal of Chemical Physics*, 1989, **90**, 1007–1023.
 - 54 A. D. Becke, *The Journal of Chemical Physics*, 1993, **98**, 5648–5652.
 - 55 P. J. Stephens, F. J. Devlin, C. F. Chabalowski and M. J. Frisch, *The Journal of Physical Chemistry*, 1994, **98**, 11623–11627.
 - 56 C. Møller and M. S. Plesset, *Phys. Rev.*, 1934, **46**, 618–622.
 - 57 K. Raghavachari, G. W. Trucks, J. A. Pople and M. Head-Gordon, *Chemical Physics Letters*, 1989, **157**, 479 – 483.
 - 58 T. H. D. Jr., K. A. Peterson and A. K. Wilson, *The Journal of Chemical Physics*, 2001, **114**, 9244–9253.
 - 59 J. F. Stanton and J. Gauss, *The Journal of Chemical Physics*, 1994, **101**, 8938–8944.
 - 60 M. Nooijen and R. J. Bartlett, *The Journal of Chemical Physics*, 1995, **102**, 3629–3647.
 - 61 D. Sinha, S. K. Mukhopadhyay, M. Prasad and D. Mukherjee, *Chemical Physics Letters*, 1986, **125**, 213 – 217.
 - 62 J. F. Stanton and R. J. Bartlett, *The Journal of Chemical Physics*, 1993, **98**, 7029–7039.
 - 63 A. I. Krylov, *Annual Review of Physical Chemistry*, 2008, **59**, 433–462.
 - 64 F. Weigend, *Phys. Chem. Chem. Phys.*, 2006, **8**, 1057–1065.
 - 65 F. Weigend and R. Ahlrichs, *Phys. Chem. Chem. Phys.*, 2005, **7**, 3297–3305.
 - 66 R. C. Fortenberry and T. D. Crawford, *The Journal of Chemical Physics*, 2011, **134**, 154304.
 - 67 R. C. Fortenberry and T. D. Crawford, *The Journal of Physical Chemistry A*, 2011, **115**, 8119–8124.
 - 68 W. J. Morgan and R. C. Fortenberry, *Theoretical Chemistry Accounts*, 2015, **134**, 47.
 - 69 W. J. Morgan and R. C. Fortenberry, *The Journal of Physical Chemistry A*, 2015, **119**, 7013–7025.
 - 70 M. K. Bassett and R. C. Fortenberry, *The Journal of Chemical Physics*, 2017, **146**, 224303.
 - 71 T. J. Santaloci and R. C. Fortenberry, *Molecular Astrophysics*, 2020, **19**, 100070.
 - 72 W. Humphrey, A. Dalke and K. Schulten, *Journal of Molecular Graphics*, 1996, **14**, 33 – 38.
 - 73 R. D. Brown, P. D. Godfrey, R. Champion and D. McNaughton, *Journal of the American Chemical Society*, 1981, **103**, 5711–5715.
 - 74 R. A. Loomis, B. A. McGuire, C. Shingledecker, C. H. Johnson, S. Blair, A. Robertson and A. J. Remijan, *The Astrophysical Journal*, 2015, **799**, 34.
 - 75 R. D. Brown, R. F. C. Brown, F. W. Eastwood, P. D. Godfrey and D. McNaughton, *Journal of the American Chemical Society*, 1979, **101**, 4705–4708.
 - 76 A. Scott and L. Radom, *Journal of Molecular Structure*, 2000, **556**, 253 – 261.
 - 77 R. D. Brown and R. G. Dittman, *Chemical Physics*, 1984, **83**, 77 – 82.
 - 78 L. Farnell and L. Radom, *Chemical Physics Letters*, 1982, **91**, 373 – 377.
 - 79 S. R. Langhoff and D. P. Chong, *Chemical Physics Letters*, 1983,

- 100, 259 – 262.
- 80 R. D. Brown, in *Structural Information on Large Amplitude Motions*, ed. J. Laane, M. Dakkouri, B. van der Veken and H. Oberhammer, Springer Netherlands, Dordrecht, 1993, pp. 99–112.
 - 81 S. Şahin, E. A. Bleda, Z. Altun and C. Trindle, *International Journal of Quantum Chemistry*, 2016, **116**, 444–451.
 - 82 K. Watanabe, T. Nakayama and J. Mottl, *Journal of Quantitative Spectroscopy and Radiative Transfer*, 1962, **2**, 369 – 382.
 - 83 J. Pitarch-Ruiz, A. Sánchez de Merás, J. Sánchez-Marán, A. M. Velasco, C. Laván and I. Martán, *The Journal of Physical Chemistry A*, 2008, **112**, 3275–3280.
 - 84 D. Truch, D. Salomon and D. Armstrong, *Journal of Molecular Spectroscopy*, 1979, **78**, 31 – 39.
 - 85 T. J. Lee, J. E. Rice, G. E. Scuseria and H. F. Schaefer, *Theoretica chimica acta*, 1989, **75**, 81–98.
 - 86 D. C. Moule and A. D. Walsh, *Chemical Reviews*, 1975, **75**, 67–84.
 - 87 M. Mühlhäuser and M. Gruber-Stadler, *Chemical Physics*, 2008, **343**, 311 – 318.
 - 88 H. G. Libuda, F. Zabel, E. H. Fink and K. H. Becker, *The Journal of Physical Chemistry*, 1990, **94**, 5860–5865.